

RAMSES use case : self-gravitating astrophysical plasma flows



BOURNAUD Frédéric – frederic.bournaud@cea.fr
CHAPON Damien - damien.chapon@cea.fr
STRAFELLA Loïc - loic.strafella@cea.fr

RAMSES ^(a) code

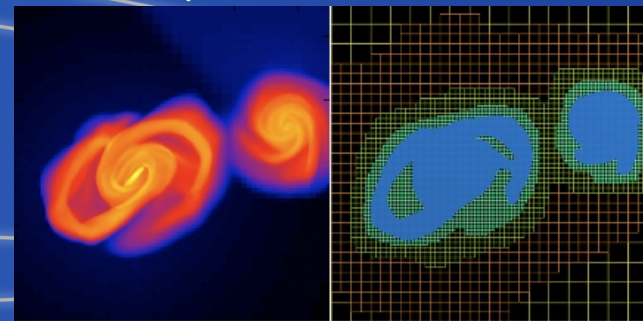
- Developed since 2002,
- Open source code to model self-gravitating astrophysical magnetised fluids,
- Community of ~ 300 users all around the world. One of the 4 most used code in astrophysics,
- Fortran 90 standard, use OpenMP and MPI, compilation with Intel and GNU compilers,

^(a) R. Teyssier, 2002, *Cosmological hydrodynamics with adaptive mesh refinement: a new high-resolution code called RAMSES*, A&A, 385, 337-364

RAMSES : introduction

RAMSES ^(a) code

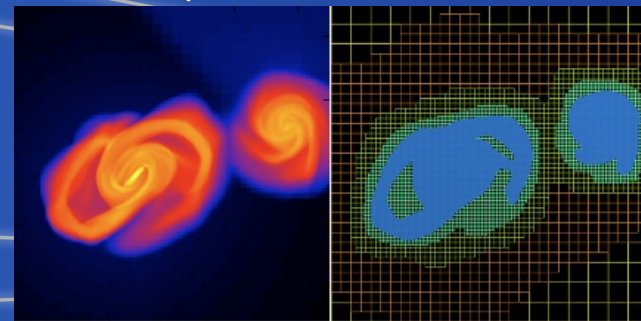
- Developed since 2002,
- Open source code to model self-gravitating astrophysical magnetised fluids,
- Community of ~ 300 users all around the world. One of the 4 most used code in astrophysics,
- Fortran 90 standard, use OpenMP and MPI, compilation with Intel and GNU compilers,
- About 50k lines of codes, adaptive mesh refinement (AMR) technique



(a) R. Teyssier, 2002, Cosmological hydrodynamics with adaptive mesh refinement: a new high-resolution code called RAMSES, A&A, 385, 337-364

RAMSES ^(a) code

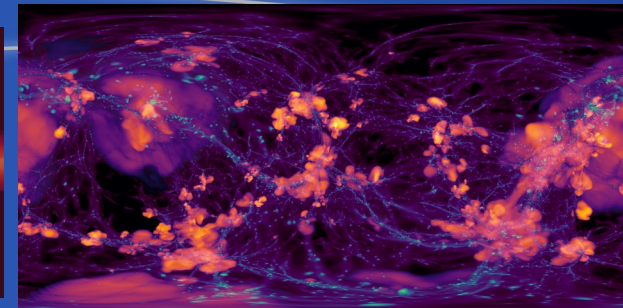
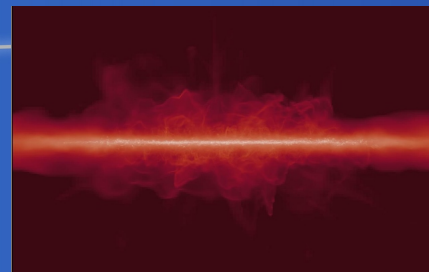
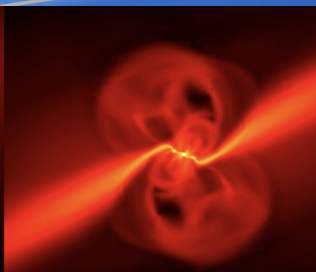
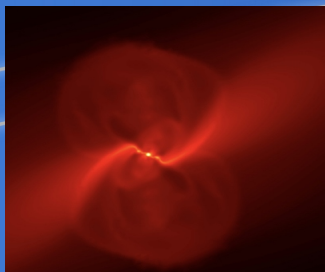
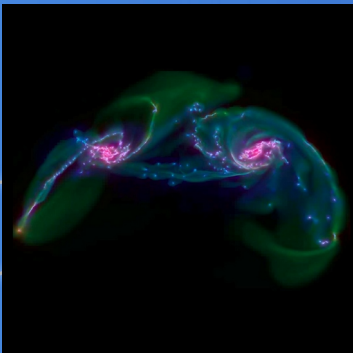
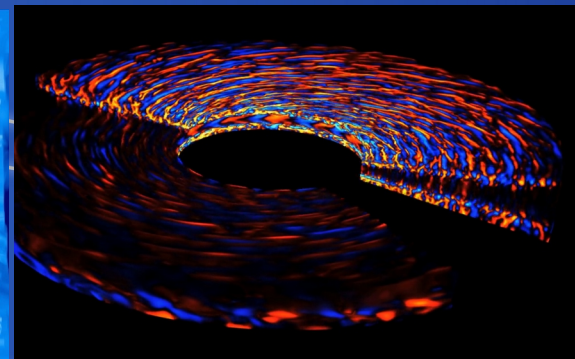
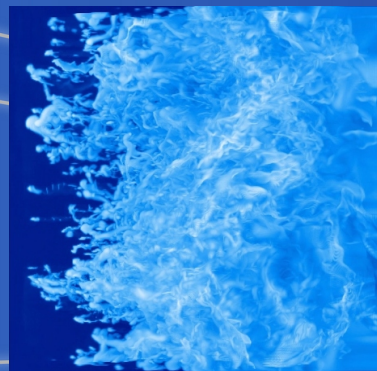
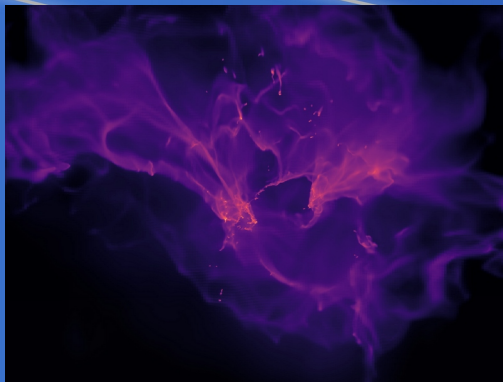
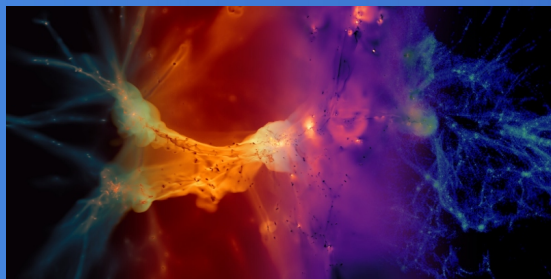
- Developed since 2002,
- Open source code to model self-gravitating astrophysical magnetised fluids,
- Community of ~ 300 users all around the world. One of the 4 most used code in astrophysics,
- Fortran 90 standard, use OpenMP and MPI, compilation with Intel and GNU compilers,
- About 50k lines of codes, adaptive mesh refinement (AMR) technique
- **I/O engine** : custom POSIX files
 - Checkpoints only, used for restarting the code and for analysis
 - Code-specific Fortran binary files
 - N files per MPI process : 10k cores simulation → **40000 – 80000 files / snapshot**

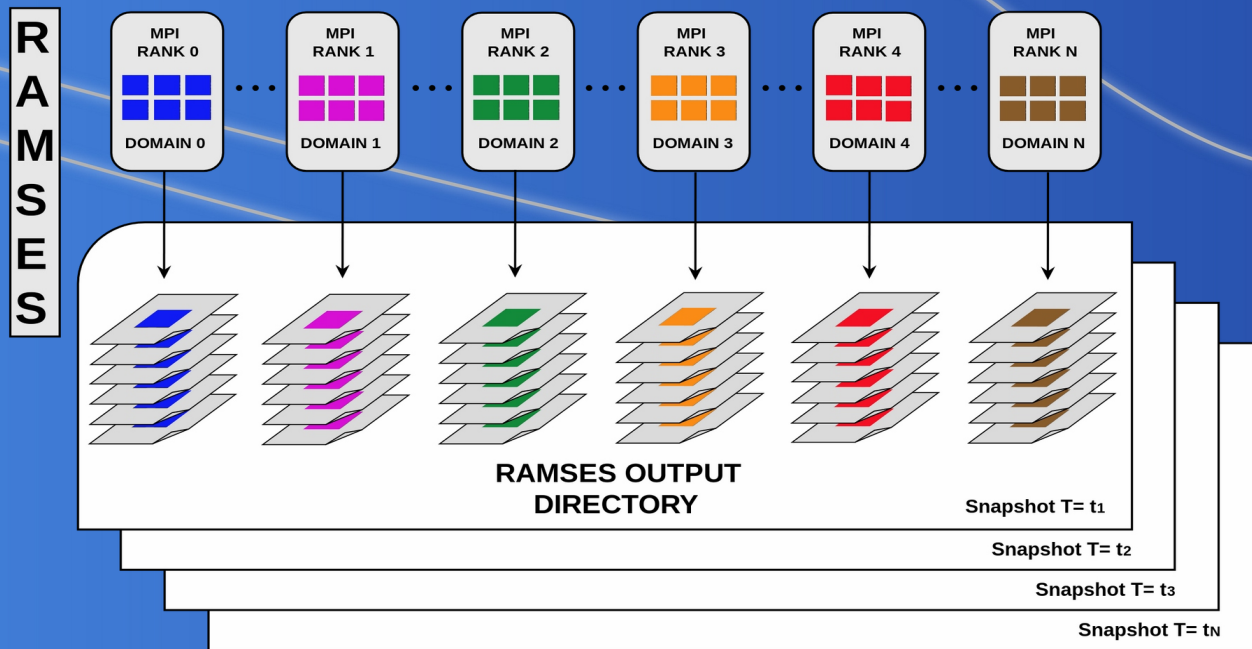


(a) R. Teyssier, 2002, Cosmological hydrodynamics with adaptive mesh refinement: a new high-resolution code called RAMSES, A&A, 385, 337-364

RAMSES is used in

- Cosmology, Galactic dynamics, Galaxy interactions,
- Interstellar medium modeling, stellar formation, planetary system formation, accretion disks,
- Supernovae explosions,





1 MPI process = n files (depending on the activated physics modules)

1 checkpoint = 1 directory = $P_{\text{process}} * n$ files

1 simulation = N snapshots = N directories = $N * P_{\text{process}} * n$ files (small files < 1GB)

➡ **RAMSES limit of scalability : ~ 8,000 MPI processes**

IO-SEA : expected benefits

Main objectives:

- ❑ Improve RAMSES I/O scalability,
- ❑ Improve post-processing productivity and data management



Main objectives:

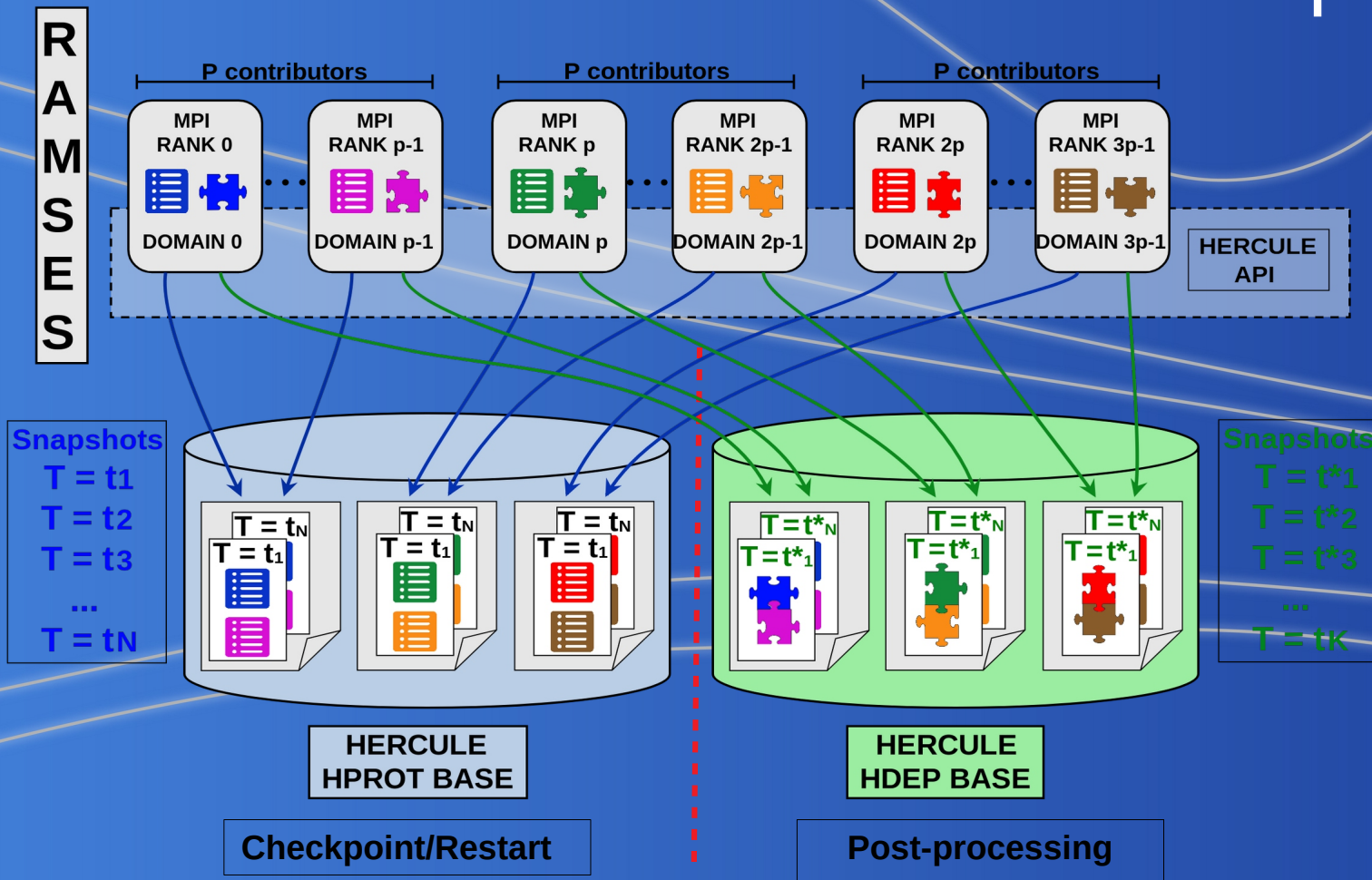
- ❑ Improve RAMSES I/O scalability,
- ❑ Improve post-processing productivity and data management



Integration of Hercule (WP5) parallel I/O library in RAMSES :

- ❑ Reduced number of files (inode limitations on Lustre filesystem)
 - contribution of multiple MPI processes to a single shared file (I/O concurrency)
- ❑ Improved scalability of RAMSES I/Os over 8,000 MPI processes,
- ❑ Taking advantage of supercomputer storage architecture
 - Multiple OSS, ...
- ❑ Semantic approach : compatible with DASI interface,
- ❑ Reduce data volume and improve simulation data management

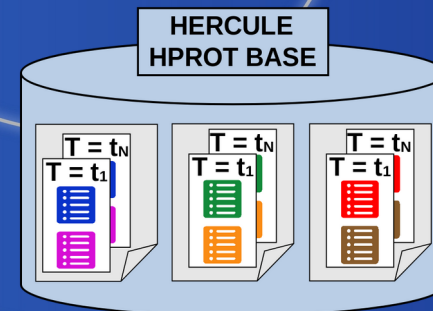
New HERCULE I/O pattern



RAMSES : dataflow separation

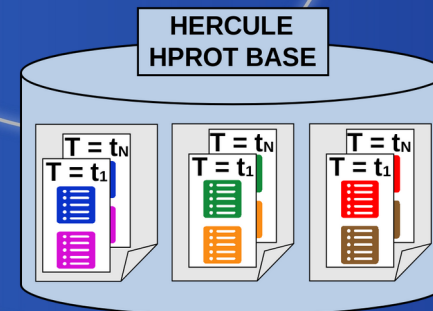
HProt database

- Dedicated exclusively to **checkpoints/restarts**
- Provide a basic binary storage
 - Understandable only by the code that produced it
- Takes advantage of asynchronous I/O (No copy of data)
 - No additional memory cost



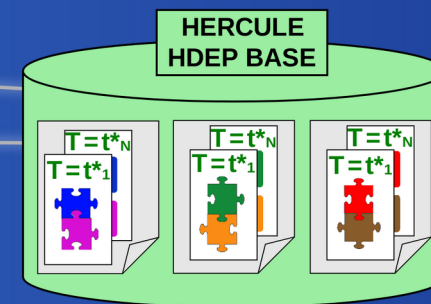
HProt database

- Dedicated exclusively to **checkpoints/restarts**
- Provide a basic binary storage
 - Understandable only by the code that produced it
- Takes advantage of asynchronous I/O (No copy of data)
 - No additional memory cost

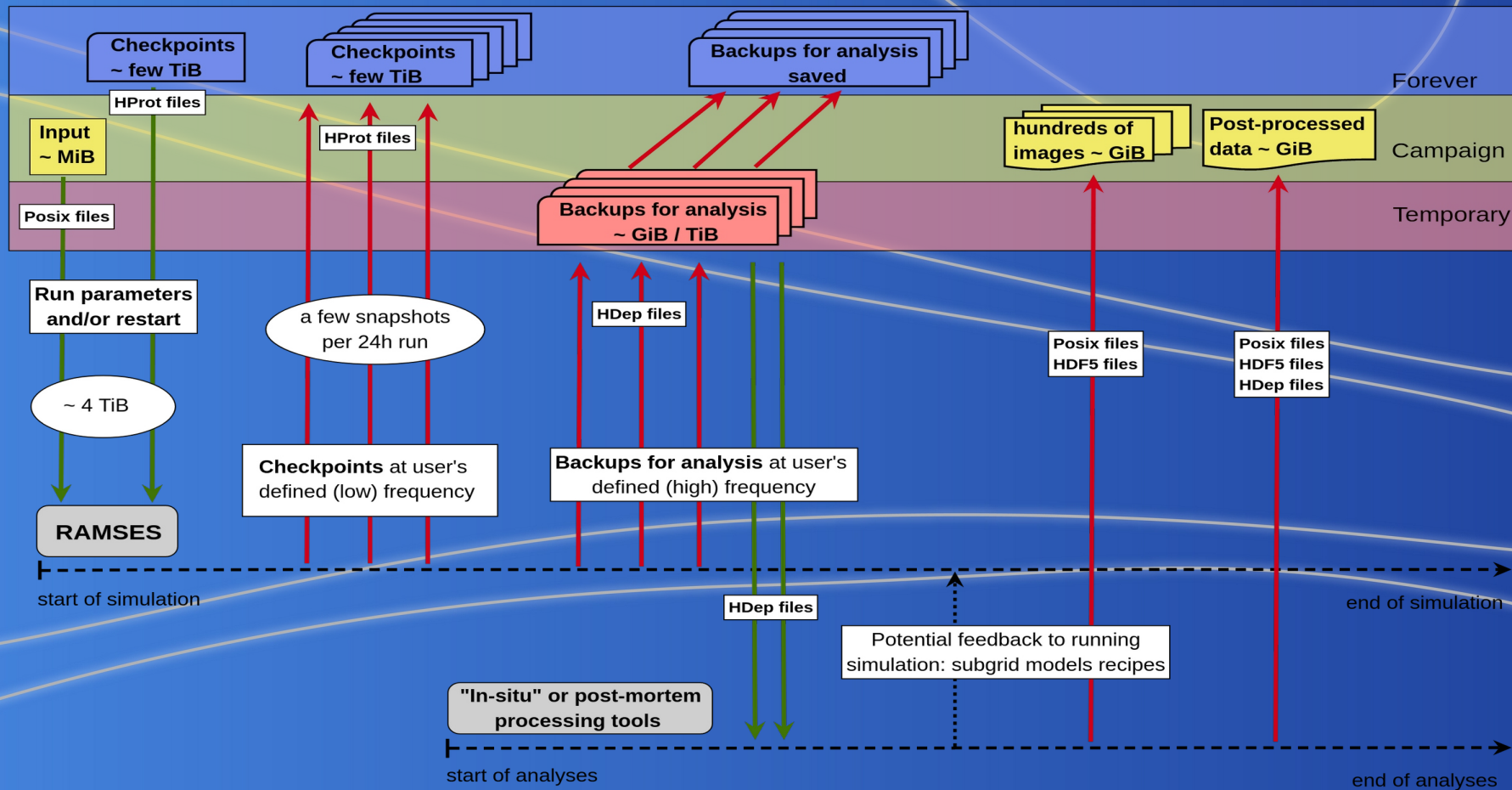


HDep database

- Dedicated exclusively to **post-processing** (analysis, visualization)
- Provide optimized data models
 - Self-described and understandable by other (post-processing codes)
- Possible restriction to a subset of the physical quantities
- Lightweight (compare to HProt and RAMSES legacy)
 - Data redundancy removal + data compression



RAMSES : workflow diagram



NFS interface

- Read initial conditions files (few MB), once per simulation,
- Read configuration files (few kB), at each step of the simulation run session,
- Write output log files (few MB).

NFS interface

- Read initial conditions files (few MB), once per simulation,
- Read configuration files (few kB), at each step of the simulation run session,
- Write output log files (few MB).

DASI interface (I/O intensive)

- Run session
 - Read/write checkpoint Hercule database (Hprot) : read upon restart, write upon checkpoint.
 - Write analysis Hercule databases (Hdep) : light-weight, but at high frequency.
- Analysis session
 - Read analysis Hercule databases (Hdep).

```
# start a session for my workflow (described in the ramses_workflow.yaml file)
ioseawf start WORKFLOW=ramses_workflow.yaml SESSION=$session

# run the workflow steps => first launch run from initial conditions
ioseawf run SESSION=$session STEP=init NPROCS=$NCORES

# Restart simulation until final completion
for i in $(seq 1 $NRESTARTS) ; do
    ioseawf run SESSION=$session STEP=restart NPROCS=$NCORES
done

# check status of jobs
ioseawf status SESSION=$session

# stop session
ioseawf stop SESSION=$session
```

Thank you for your attention.
Questions ?